# Theoretical optical potential derived from nucleon-nucleon chiral potentials 

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#### Abstract

Background: Elastic scattering is probably the main event in the interactions of nucleons with nuclei. Even if this process has been extensively studied over the last years, a consistent description, i.e., starting from microscopic two- and many-body forces connected by the same symmetries and principles, is still under development. Purpose: In this work we study the domain of applicability of microscopic two-body chiral potentials in the construction of an optical potential. Methods: We basically follow the Kerman, McManus, and Thaler approach [Ann. Phys. (NY) 8, 551 (1959)] to build a microscopic complex optical potential, and then we perform some test calculations on ${ }^{16} \mathrm{O}$ at different energies. Results:. Our conclusion is that a particular set of potentials with a Lippmann-Schwinger cutoff at relatively high energies (above 500 MeV ) reproduces best the scattering observables. Conclusions: Our work shows that building an optical potential within chiral perturbation theory is a promising approach for describing elastic proton scattering; in particular, in view of the future inclusion of many-body forces that naturally arises in such a framework.


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## I. INTRODUCTION

Elastic proton scattering has been extensively studied over many decades, both experimentally and theoretically, and there now exist extensive measurements of cross sections and polarization observables for the elastic scattering of protons from a wide variety of stable nuclei over a wide range of energies. A suitable and successful framework to describe the nucleon-nucleus ( $N A$ ) interaction in elastic scattering is provided by the nuclear optical potential [1]. With this instrument we can compute the scattering observables, such as the differential cross section, the analyzing power, and the spin rotation, for the elastic $N A$ scattering across wide regions of the nuclear landscape. The use of the optical potential has been extended to calculations of inelastic scattering and to generate the distorted waves for the analysis of the cross sections for a wide variety of nuclear reactions. For instance, in quasi-elastic electron scattering, an optical potential is commonly used to describe the final-state interaction between the emitted nucleon and the residual nucleus in the exclusive $\left(e, e^{\prime} p\right)$ [2] and in the inclusive ( $e, e^{\prime}$ ) reactions [3,4].

The optical potential can be obtained in different ways. It can be obtained phenomenologically [5,6], by assuming a form of the potential and a dependence on a number of adjustable parameters for the real and imaginary parts that characterizes the shape of the nuclear density distribution and that varies with the nucleon energy and with the nucleus mass number. These parameters are adjusted to optimize the fit to the experimental data of elastic $N A$ scattering. The optical potential has an imaginary part that takes into account the absorption of the reaction flux from the elastic channel to the nonelastic reaction channels. Alternatively and more fundamentally, the optical potential can be obtained microscopically. The calculation requires, in principle, the solution of the full nuclear many-body problem, which is beyond present capabilities. In practice, some approximations must necessarily be adopted. With suitable approximations,
microscopic optical potentials are usually derived from two basic quantities: the nucleon-nucleon $(N N) t$ matrix and the matter distribution of the nucleus. All these models based on the $N N$ interaction are nonrelativistic (see Ref. [7] for a detailed review). Because microscopic optical potentials do not contain adjustable parameters, we expect that they have a greater predictive power when applied to situations where experimental data are not yet available, such as, for instance, to the study of unstable nuclei.

The theoretical justification for the $N A$ optical potential built in terms of underlying $N N$ scattering amplitudes was given for the first time by Chew [8] and Watson et al. [9,10] more than 60 years ago. Successively, Kerman, McManus, and Thaler (KMT) [11] developed the Watson multiple scattering approach expressing the $N A$ optical potential by a series expansion in terms of the free $N N$ scattering amplitudes. Several years later, with the development of high-accuracy $N N$ potentials, there has been a renewed interest in finding a rigorous treatment of the $N A$ scattering theory in momentum space. Such potentials permit us to generate the $N N$ interaction directly in momentum space, which is thus chosen as the working space in which to develop the $N A$ optical potential and to compute the elastic scattering observables. Several authors contributed to the development of multiple scattering theory and, with a series of papers [12-32], to calculations of microscopic optical potentials. The present work is framed in this context.

The $N N$ potential is an essential ingredient in the $N A$ scattering theory and its off-shell properties play an important role. To obtain a good description of these properties, the optical potential models have always employed "realistic" potentials, in which the experimental $N N$ phase shifts are reproduced with a $\chi^{2}$ per data $\simeq 1$. The most commonly used $N N$ potentials are those given by groups from Nijmegen [33], Paris [34], Bonn [35], and Argonne [36]. In contrast, with recent advances in lattice quantum chromodynamics
(lattice QCD), new attempts have been made to derive the nuclear potential directly from first principles [37]. However, since QCD is a nonperturbative theory in the low-energy regime, which is characteristic of nuclear physics, an ab initio solution of this problem is not feasible at the moment. On the other hand, when the concepts of effective field theory (EFT) were applied to low-energy QCD, chiral perturbation theory (ChPT) [38] was developed and it became possible to implement chiral symmetry consistently in a theory of pionic and nuclear interactions. For this theory some sort of perturbative expansion was assumed, such that only a finite number of terms contribute at a given order. This expansion was provided by powers of small external momenta over the chiral-symmetry-breaking scale, $\Lambda_{\chi} \approx 1 \mathrm{GeV}$. The $N N$ potential is then calculated perturbatively in the chiral expansion and iterated to all orders in a Schrödinger or Lippmann-Schwinger (LS) equation to obtain the nuclear amplitude.

The most recently available chiral potentials are developed at fourth order $\left(\mathrm{N}^{3} \mathrm{LO}\right)$ in the chiral expansion and are used in this work as a basic ingredient to compute the $N N t$ matrix for the construction of the $N A$ optical potential. In particular, in all the calculations presented in this paper we adopt the two different versions of chiral potentials developed by Entem and Machleidt (EM, [39-43]), and Epelbaum, Glöckle, and Meißner (EGM, [44]).

Very recently Epelbaum et al. [45] presented a nucleonnucleon potential at fifth order ( $\mathrm{N}^{4} \mathrm{LO}$ ) with an innovative coordinate-space regularization. It is also worth mentioning that, over the last years, some authors [46-48], following different approaches, have started to include chiral threebody forces $[49,50]$ at order $\mathrm{N}^{2} \mathrm{LO}$. We plan to extend our calculations along these research lines in forthcoming presentations.

The second important ingredient of the $N A$ scattering theory is the microscopic structure of the nuclear target, given by neutron and proton densities. These quantities are computed within the relativistic mean-field (RMF) description [51] of spherical nuclei by using a density-dependent mesonexchange (DDME) model, where the couplings between mesonic and baryonic fields are assumed as functions of the density itself [52].

The paper is organized as follows: in Sec. II we describe the theoretical framework used to calculate the $N A$ optical potential and the scattering observables. This section is divided into five sections in which we outline the different aspects of the calculation. In Sec. II A the general scattering problem is introduced in the momentum frame and is represented by the LS equation for the entire system composed of the projectile and the target nucleus. This equation is separated into a simple one-body equation for the transition matrix and a more complicated one for the optical potential. By using the optimum factorization approximation the expression for the optical potential is reduced to a simple form, in which the $N N t$ matrix and the nuclear density are factorized. In Sec. II B we give the explicit formulas to compute the $N N$ Wolfenstein amplitudes that are proportional to the central and spin-orbit parts of the $N N t$ matrix, which is then used to compute the $N A$ optical potential. In Secs. II C and II D we
describe the theoretical framework to solve the $N A$ LS equation in the partial-wave representation and then use these solutions to compute the scattering observables. In Sec. IIE we show the algorithm we use to include in the model the Coulomb interaction between the projectile and the target nucleus.

In Sec. III we present and discuss the theoretical results for the $N N$ Wolfenstein amplitudes obtained with the different $N N$ potentials. In particular, the novelty in our calculations is the use of the chiral potentials $[39,44]$ as basic ingredients to compute the microscopic $N A$ optical potential and the scattering observables.

In Sec. IV we present theoretical results for the scattering observables on ${ }^{16} \mathrm{O}$ calculated with all $N N$ potentials. Predictions based on EM and EGM potentials are compared with available experimental data in order to determine the most successful theoretical approach and the best LS cutoff.

Finally, in Sec. V, we draw our conclusions.

## II. THEORETICAL FRAMEWORK

## A. First-order optical potential

The general problem of the elastic scattering of a proton from a target nucleus of $A$ nucleons can be stated in momentum space by the full $(A+1)$-body Lippmann-Schwinger equation

$$
\begin{equation*}
T=V+V G_{0}(E) T \tag{1}
\end{equation*}
$$

whose general solution is beyond present capabilities. A reliable method to treat Eq. (1) is given by the spectator expansion [30], in which multiple scattering theory is expanded in a finite series of terms where the target nucleons interact directly with the incident proton. In particular, the first term of this series only involves the interaction of the projectile with a single target nucleon, the second term involves the interaction of the projectile with two target nucleons, and so on to the subsequent orders. In the standard approach to elastic scattering, Eq. (1) is separated into two equations. The first one is an integral equation for $T$

$$
\begin{equation*}
T=U+U G_{0}(E) P T \tag{2}
\end{equation*}
$$

where $U$ is the optical potential operator, and the second one is an integral equation for $U$

$$
\begin{equation*}
U=V+V G_{0}(E) Q U \tag{3}
\end{equation*}
$$

The operator $V$ represents the external interaction and the total Hamiltonian for the $(A+1)$-nucleon system is given by

$$
\begin{equation*}
H_{A+1}=H_{0}+V \tag{4}
\end{equation*}
$$

If we assume the presence of only two-body forces, the operator $V$ is expressed as

$$
\begin{equation*}
V=\sum_{i=1}^{A} v_{0 i} \tag{5}
\end{equation*}
$$

where the two-body potential $v_{0 i}$ describes the interaction between the incident proton and the $i$ th target nucleon. The system is asymptotically an eigenstate of the free Hamiltonian $H_{0}$ and $G_{0}(E)$ is the free propagator for the $(A+1)$-nucleon
system,

$$
\begin{equation*}
G_{0}(E)=\frac{1}{E-H_{0}+i \epsilon} \tag{6}
\end{equation*}
$$

The free Hamiltonian is given by

$$
\begin{equation*}
H_{0}=h_{0}+H_{A}, \tag{7}
\end{equation*}
$$

where $h_{0}$ is the kinetic energy operator of the projectile and $H_{A}$ is the target Hamiltonian,

$$
\begin{equation*}
H_{A}\left|\Phi_{A}\right\rangle=E_{A}\left|\Phi_{A}\right\rangle \tag{8}
\end{equation*}
$$

where $\left|\Phi_{A}\right\rangle$ is the ground state of the target. The operators $P$ and $Q$ in Eqs. (2) and (3) are projection operators,

$$
\begin{equation*}
P+Q=\mathbb{1} \tag{9}
\end{equation*}
$$

and $P$ fulfills the condition

$$
\begin{equation*}
\left[G_{0}, P\right]=0 \tag{10}
\end{equation*}
$$

In the case of elastic scattering $P$ projects onto the elastic channel and can be defined as

$$
\begin{equation*}
P=\frac{\left|\Phi_{A}\right\rangle\left\langle\Phi_{A}\right|}{\left\langle\Phi_{A} \mid \Phi_{A}\right\rangle} . \tag{11}
\end{equation*}
$$

With these definitions, the elastic transition operator may be defined as $T_{\mathrm{el}}=P T P$ and, in this case, Eq. (2) becomes

$$
\begin{equation*}
T_{\mathrm{el}}=P U P+P U P G_{0}(E) T_{\mathrm{el}} . \tag{12}
\end{equation*}
$$

Thus the transition operator for elastic scattering is given by a one-body integral equation. In order to solve Eq. (12) we need to know the operator $P U P$. In the spectator expansion the operator $U$ is expanded as

$$
\begin{equation*}
U=\sum_{i=1}^{A} \tau_{i}+\sum_{i, j \neq i}^{A} \tau_{i j}+\sum_{i, j \neq i, k \neq i, j}^{A} \tau_{i j k}+\cdots \tag{13}
\end{equation*}
$$

according to the number of nucleons interacting with the projectile. In the present work we only consider the first-order term of this expansion and thus the optical potential operator becomes

$$
\begin{equation*}
U=\sum_{i=1}^{A} \tau_{i} \tag{14}
\end{equation*}
$$

where $\tau_{i}$ can be expressed as

$$
\begin{equation*}
\tau_{i}=\hat{\tau}_{i}-\hat{\tau}_{i} G_{0}(E) P \tau_{i} . \tag{15}
\end{equation*}
$$

For elastic scattering we only need to consider $P \tau_{i} P$ or, equivalently,

$$
\begin{align*}
\left\langle\Phi_{A}\right| \tau_{i}\left|\Phi_{A}\right\rangle= & \left\langle\Phi_{A}\right| \hat{\tau}_{i}\left|\Phi_{A}\right\rangle-\left\langle\Phi_{A}\right| \hat{\tau}_{i}\left|\Phi_{A}\right\rangle \\
& \times \frac{1}{\left(E-E_{A}\right)-h_{0}+i \epsilon}\left\langle\Phi_{A}\right| \tau_{i}\left|\Phi_{A}\right\rangle, \tag{16}
\end{align*}
$$

where $\hat{\tau}_{i}$ is the solution of

$$
\begin{equation*}
\hat{\tau}_{i}=v_{0 i}+v_{0 i} G_{0}(E) \hat{\tau}_{i} . \tag{17}
\end{equation*}
$$

Expanding the propagator $G_{0}(E)$ within a single-particle description, at first order we obtain

$$
\begin{equation*}
G_{i}(E)=\frac{1}{\left(E-E^{i}\right)-h_{0}-h_{i}-W_{i}+i \epsilon} \tag{18}
\end{equation*}
$$

where $h_{i}$ is the kinetic energy of the $i$ th target nucleon and $W_{i}$ is given by

$$
\begin{equation*}
W_{i}=\sum_{j \neq i} v_{i j} \tag{19}
\end{equation*}
$$

and represents the force between the struck nucleon and the other $(A-1)$ nucleons. With the operator $G_{i}(E)$, the $\hat{\tau}_{i}$ matrix of Eq. (17) is expressed as

$$
\begin{equation*}
\hat{\tau}_{i}=v_{0 i}+v_{0 i} G_{i}(E) \hat{\tau}_{i}=t_{0 i}+t_{0 i} g_{i} W_{i} G_{i}(E) \hat{\tau}_{i}, \tag{20}
\end{equation*}
$$

where the operators $t_{0 i}$ and $g_{i}$ are defined as

$$
\begin{align*}
t_{0 i} & =v_{0 i}+v_{0 i} g_{i} t_{0 i}  \tag{21}\\
g_{i} & =\frac{1}{\left(E-E^{i}\right)-h_{0}-h_{i}+i \epsilon} \tag{22}
\end{align*}
$$

In Eq. (21) the matrix $t_{0 i}$ represents the free $N N t$ matrix and, in the impulse approximation (IA), we have $\hat{\tau}_{i} \approx t_{0 i}$. Thus in this approximation we only have to solve a two-body equation.

To develop a theoretical framework to compute the optical potential and the transition amplitude for the elastic scattering observables, we follow the path outlined in Ref. [53], that is based on the KMT multiple scattering theory and that, at first order, is equivalent to the IA. In this formulation, the elastic scattering amplitude is given by

$$
\begin{equation*}
T_{\mathrm{el}}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right)=\frac{A}{A-1} \hat{T}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right) \tag{23}
\end{equation*}
$$

where the auxiliary elastic amplitude is determined by the solution of the integral equation

$$
\begin{equation*}
\hat{T}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right)=\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)+\int d^{3} p \frac{\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{p} ; \omega\right) \hat{T}(\boldsymbol{p}, \boldsymbol{k} ; E)}{E\left(k_{0}\right)-E(p)+i \epsilon}, \tag{24}
\end{equation*}
$$

and the auxiliary first-order optical potential is defined by

$$
\begin{equation*}
\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)=(A-1)\left\langle\boldsymbol{k}^{\prime}, \Phi_{A}\right| t(\omega)\left|\boldsymbol{k}, \Phi_{A}\right\rangle, \tag{25}
\end{equation*}
$$

where $t$ is any one of the free $N N t_{0 i}$ matrices. Our problem is then described in the zero-momentum frame of the $N A$ system by Eq. (24), where $k_{0}$ is the initial on-shell momentum and $E\left(k_{0}\right)$ is the corresponding initial energy of the system in the $N A$ frame. To compute the scattering observables we only need the on-shell term $T_{\mathrm{el}}\left(\boldsymbol{k}_{0}, \boldsymbol{k}_{0} ; E\right)$ of the transition matrix, but in this work we consider the full off-shell matrix with the general initial and final momenta $\boldsymbol{k}$ and $\boldsymbol{k}^{\prime}$, respectively.

The KMT first-order optical potential is given by Eq. (25) where $t(\omega)$ is the free $N N t$ matrix evaluated at a fixed energy $\omega$. Defining the new variables

$$
\begin{equation*}
\boldsymbol{q} \equiv \boldsymbol{k}^{\prime}-\boldsymbol{k}, \quad \boldsymbol{K} \equiv \frac{1}{2}\left(\boldsymbol{k}^{\prime}+\boldsymbol{k}\right) \tag{26}
\end{equation*}
$$

some manipulations [53] give Eq. (25) in a factorized form (optimum factorization approximation) as the product of the $N N t$ matrix and the nuclear matter density,
$\hat{U}(\boldsymbol{q}, \boldsymbol{K} ; \omega)=\frac{A-1}{A} \eta(\boldsymbol{q}, \boldsymbol{K}) \sum_{N=n, p} t_{p N}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right] \rho_{N}(q)$,
where $N=n, p, t_{p N}$ represents the proton-proton ( pp ) and proton-neutron (pn) $t$ matrix, $\rho_{N}$ the neutron and proton profile
density, and $\eta(\boldsymbol{q}, \boldsymbol{K})$ is the Møller factor,
$\eta(\boldsymbol{q}, \boldsymbol{K})=\left[\frac{E_{\text {proj }}\left(\boldsymbol{\kappa}^{\prime}\right) E_{\mathrm{proj}}\left(-\boldsymbol{\kappa}^{\prime}\right) E_{\mathrm{proj}}(\boldsymbol{\kappa}) E_{\mathrm{proj}}(-\boldsymbol{\kappa})}{E_{\mathrm{proj}}\left(\boldsymbol{k}^{\prime}\right) E_{\mathrm{proj}}\left(-\frac{q}{2}-\frac{\boldsymbol{K}}{A}\right) E_{\mathrm{proj}}(\boldsymbol{k}) E_{\mathrm{proj}}\left(\frac{q}{2}-\frac{K}{A}\right)}\right]^{\frac{1}{2}}$,
that imposes the Lorentz invariance of the flux when we pass from the $N A$ to the $N N$ frame in which the $t$ matrices are evaluated. The optical potential obtained so far is an operator in the spin space of the projectile. To make the spin dependence explicit, the $t$ matrix $t_{p N}$ is averaged over the spin of the struck nucleon and is written as

$$
\begin{align*}
t_{p N}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right]= & t_{p N}^{c}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right]+\left(\frac{A+1}{2 A}\right) \\
& \times \frac{i}{2} \boldsymbol{\sigma} \cdot \boldsymbol{q} \times \boldsymbol{K} t_{p N}^{l s}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right] . \tag{29}
\end{align*}
$$

The first term of Eq. (29) corresponds to the central spinindependent contribution and the second term corresponds to the spin-orbit contribution. In the latter term the usual total Pauli spin operator of the $N N$ system is replaced by the Pauli spin operator of the projectile, because the spin operator of the struck nucleon has been eliminated by the trace over the spin. The replacement of Eq. (29) into Eq. (27) gives the optical potential as

$$
\begin{equation*}
\hat{U}(\boldsymbol{q}, \boldsymbol{K} ; \omega)=\hat{U}^{c}(\boldsymbol{q}, \boldsymbol{K} ; \omega)+\frac{i}{2} \boldsymbol{\sigma} \cdot \boldsymbol{q} \times \boldsymbol{K} \hat{U}^{l s}(\boldsymbol{q}, \boldsymbol{K} ; \omega) \tag{30}
\end{equation*}
$$

where the central and the spin-orbit terms are given by

$$
\begin{align*}
\hat{U}^{c}(\boldsymbol{q}, \boldsymbol{K} ; \omega)= & \frac{A-1}{A} \eta(\boldsymbol{q}, \boldsymbol{K}) \\
& \times \sum_{N=n, p} t_{p N}^{c}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right] \rho_{N}(q)  \tag{31}\\
\hat{U}^{l s}(\boldsymbol{q}, \boldsymbol{K} ; \omega)= & \frac{A-1}{A} \eta(\boldsymbol{q}, \boldsymbol{K})\left(\frac{A+1}{2 A}\right) \\
& \times \sum_{N=n, p} t_{p N}^{l s}\left[\boldsymbol{q}, \frac{A+1}{A} \boldsymbol{K} ; \omega\right] \rho_{N}(q) \tag{32}
\end{align*}
$$

The optimally factorized optical potential given in Eqs. (30)-(32) exhibits nonlocality and off-shell effects through the dependence of $\eta$ and $t_{p N}$ upon $\boldsymbol{K}$. The energy $\omega$ at which the matrices $t_{p N}^{c}$ and $t_{p N}^{l s}$ are evaluated is fixed as

$$
\begin{equation*}
\omega=\frac{T_{\mathrm{lab}}}{2}=\frac{1}{2} \frac{k_{\mathrm{lab}}^{2}}{2 m} \tag{33}
\end{equation*}
$$

where $k_{\text {lab }}$ is the on-shell momentum of the projectile in the laboratory system. This is the fixed beam energy approximation, which is a historic choice performed in all calculations based on the KMT formulation. A review of this type of calculations can be found in Ref. [54].

## B. The $N N$ transition matrix

The $N N$ elastic scattering amplitude for the scattering from a relative momentum $\boldsymbol{\kappa}$ to $\boldsymbol{\kappa}^{\prime}$, denoted by $M\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa}, \omega\right)$, is related
to the antisymmetrized transition matrix elements by the usual relation $(\hbar=1)$

$$
\begin{equation*}
M\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa}, \omega\right)=\left\langle\boldsymbol{\kappa}^{\prime}\right| M(\omega)|\boldsymbol{\kappa}\rangle=-4 \pi^{2} \mu\left\langle\boldsymbol{\kappa}^{\prime}\right| t(\omega)|\boldsymbol{\kappa}\rangle \tag{34}
\end{equation*}
$$

where $\mu$ is the $N N$ reduced mass. The most general form of this amplitude, consistent with invariance under rotation, time reversal, and parity is [55]

$$
\begin{align*}
M= & a+c\left(\boldsymbol{\sigma}_{1}+\boldsymbol{\sigma}_{2}\right) \cdot \hat{\boldsymbol{n}}+m\left(\boldsymbol{\sigma}_{1} \cdot \hat{\boldsymbol{n}}\right)\left(\boldsymbol{\sigma}_{2} \cdot \hat{\boldsymbol{n}}\right)+(g+h)\left(\boldsymbol{\sigma}_{1} \cdot \hat{\boldsymbol{l}}\right) \\
& \times\left(\boldsymbol{\sigma}_{2} \cdot \hat{\boldsymbol{l}}\right)+(g-h)\left(\boldsymbol{\sigma}_{1} \cdot \hat{\boldsymbol{m}}\right)\left(\boldsymbol{\sigma}_{2} \cdot \hat{\boldsymbol{m}}\right), \tag{35}
\end{align*}
$$

where

$$
\begin{equation*}
\hat{l}=\frac{\kappa^{\prime}+\kappa}{\left|\kappa^{\prime}+\kappa\right|}, \quad \hat{\boldsymbol{m}}=\frac{\kappa^{\prime}-\kappa}{\left|\kappa^{\prime}-\kappa\right|}, \quad \hat{\boldsymbol{n}}=\frac{\kappa \times \kappa^{\prime}}{\left|\kappa \times \kappa^{\prime}\right|} \tag{36}
\end{equation*}
$$

are the unit vectors defined by the $N N$ scattering plane. The amplitudes $a, c, m, g$, and $h$ can be expressed as complex functions of $\omega, \boldsymbol{\kappa}$, and $\boldsymbol{\kappa}^{\prime}$. The amplitudes in Eq. (35) are given in the Hoshizaki notation [56]. There are different ways to define them, a survey of the other decompositions can be found in Refs. [57,58] and references therein. We may also note that, for an even-even nucleus with $J=0$, terms linear in the spin of the target nucleons average to zero; only $a$ and $c$ amplitudes survive and they are connected to the central and spin-orbit part of the $N N t$ matrix, respectively.

The $N N$ amplitudes are usually expressed in terms of the decomposition of the scattering amplitude into components describing spin singlet $(S=0)$ and spin triplet $(S=1)$ scattering, $M_{\nu^{\prime} v}^{S}$, where $v$ and $v^{\prime}$ refer to the incident- and final-spin projections in the triplet state. In the representation in which these projections are referred to an axis of quantization along the incident beam direction $(\kappa)$ we have

$$
\begin{align*}
a & =\frac{1}{4}\left(2 M_{11}^{1}+M_{00}^{1}+M_{00}^{0}\right)  \tag{37}\\
c & =\frac{i}{2 \sqrt{2}}\left(M_{10}^{1}-M_{01}^{1}\right)  \tag{38}\\
m & =\frac{1}{4}\left(M_{00}^{1}-2 M_{1-1}^{1}-M_{00}^{0}\right),  \tag{39}\\
g & =\frac{1}{4}\left(M_{11}^{1}-M_{00}^{0}+M_{1-1}^{1}\right),  \tag{40}\\
h & =\frac{1}{4 \cos \phi}\left(M_{11}^{1}-M_{00}^{1}-M_{1-1}^{1}\right) \tag{41}
\end{align*}
$$

The amplitudes $M_{v^{\prime} v}^{S}=\left\langle\boldsymbol{\kappa} S v^{\prime}\right| M(\omega)|\boldsymbol{\kappa} S v\rangle$ and hence $a-h$, are obtained [13,59] in terms of the partial-wave components of the $N N$ amplitude, $M_{L^{\prime} L}^{J S}\left(\kappa^{\prime}, \kappa ; \omega\right)$, defined by

$$
\begin{align*}
M\left(\kappa^{\prime}, \boldsymbol{\kappa} ; \omega\right)= & \frac{2}{\pi} \sum_{J L L^{\prime} S M} i^{L-L^{\prime}} y_{J M}^{L^{\prime} S}\left(\hat{\kappa}^{\prime}\right) \\
& \times M_{L^{\prime} L}^{J S}\left(\kappa^{\prime}, \kappa ; \omega\right) y_{J M}^{L S \dagger}(\hat{\kappa}) \tag{42}
\end{align*}
$$

where $y_{J M}^{L S}$ is the spin-angular function

$$
\begin{equation*}
y_{J M}^{L S}(\hat{\kappa})=\sum_{\Lambda v}(L \Lambda S v \mid J M) Y_{L}^{\Lambda}(\hat{\kappa}) \otimes \chi_{S v} \tag{43}
\end{equation*}
$$

and $Y_{L}^{\Lambda}$ and $\chi_{S \nu}$ are the spherical harmonic and the spin wave function of the $N N$ pair, respectively. Explicitly, we have

$$
\begin{align*}
M_{\nu^{\prime} \nu}^{S}= & \frac{2}{\pi} \sum_{J M L L^{\prime} \Lambda \Lambda^{\prime}} i^{L-L^{\prime}}\left(L^{\prime} \Lambda^{\prime} S \nu^{\prime} \mid J M\right)(L \Lambda S v \mid J M) \\
& \times Y_{L^{\prime}}^{\Lambda^{\prime}}\left(\hat{\kappa}^{\prime}\right) Y_{L}^{\Lambda *}(\hat{\kappa}) M_{L^{\prime} L}^{J S}\left(\kappa^{\prime}, \kappa ; \omega\right) \tag{44}
\end{align*}
$$

Detailed formulas for the required $M_{\nu^{\prime} v}^{S}$ amplitudes in terms of the partial-wave amplitudes $M_{L^{\prime} L}^{J S}\left(\kappa^{\prime}, \kappa ; \omega\right)$ for a quantization axis along the incident-beam direction can be found in Refs. [13,59]. According to these formulas, the $a$ and $c$ amplitudes are given by

$$
\begin{align*}
a_{p N}= & \frac{1}{f_{p N} \pi^{2}} \sum_{L=0}^{\infty} P_{L}(\cos \phi)\left[(2 L+1) M_{L L}^{L, S=0}\right. \\
& +(2 L+1) M_{L L}^{L, S=1}+(2 L+3) M_{L L}^{L+1, S=1} \\
& \left.+(2 L-1) M_{L L}^{L-1, S=1}\right],  \tag{45}\\
c_{p N}= & \frac{i}{f_{p N} \pi^{2}} \sum_{L=1}^{\infty} P_{L}^{1}(\cos \phi)\left[\left(\frac{2 L+3}{L+1}\right) M_{L L}^{L+1, S=1}\right. \\
& \left.-\left(\frac{2 L+1}{L(L+1)}\right) M_{L L}^{L, S=1}-\left(\frac{2 L-1}{L}\right) M_{L L}^{L-1, S=1}\right], \tag{46}
\end{align*}
$$

where $f_{p p}=4, f_{p n}=8$, and $P_{L}^{1}(x)$ are the associated Legendre polynomials

$$
\begin{equation*}
P_{L}^{1}(x)=\sqrt{1-x^{2}} \frac{d}{d x} P_{L}(x) \tag{47}
\end{equation*}
$$

From these equations we can obtain the explicit expressions for the $p p$ and $p n$ central and spin-orbit parts of the $N N t$ matrix. As stated above, the optical potential is an operator in the spin space of the projectile and the spin dependence is made explicit writing the $t$ matrix in the form $(N=p, n)$

$$
\begin{equation*}
t_{p N}\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa} ; \omega\right)=t_{p N}^{c}\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa} ; \omega\right)+\frac{i}{2} \boldsymbol{\sigma} \cdot \boldsymbol{\kappa}^{\prime} \times \boldsymbol{\kappa} t_{p N}^{l s}\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa} ; \omega\right) . \tag{48}
\end{equation*}
$$

In terms of the partial-wave components $t_{J L L}^{S T}\left(\kappa^{\prime}, \kappa ; \omega\right)$ we have the following results for the central part:

$$
\begin{align*}
t_{p p}^{c}= & \frac{1}{4 \pi^{2}} \sum_{L=0}^{\infty} P_{L}(\cos \phi)\left[(2 L+1) t_{L, L L}^{S=0, T=1}\right. \\
& +(2 L+1) t_{L, L L}^{S=1, T=1}+(2 L-1) t_{L-1, L L}^{S=1, T=1} \\
& \left.+(2 L+3) t_{L+1, L L}^{S=1, T=1}\right],  \tag{49}\\
t_{p n}^{c}= & \frac{1}{8 \pi^{2}} \sum_{L=0}^{\infty} P_{L}(\cos \phi)\left[(2 L+1) t_{L, L L}^{S=0, T=0}\right. \\
& +(2 L+1) t_{L, L L}^{S=1, T=0}+(2 L-1) t_{L-1, L L}^{S=1, T=0} \\
& +(2 L+3) t_{L+1, L L}^{S=1, T=0}+(2 L+1) t_{L, L L}^{S=0, T=1} \\
& +(2 L+1) t_{L, L L}^{S=1, T=1}+(2 L-1) t_{L-1, L L}^{S=1, T=1} \\
& \left.+(2 L+3) t_{L+1, L L}^{S=1, T=1}\right], \tag{50}
\end{align*}
$$

TABLE I. Partial waves of the $N N$ potential used to construct the three-dimensional $N N t$ matrix $t_{p N}\left(\boldsymbol{\kappa}^{\prime}, \boldsymbol{\kappa} ; \omega\right)$.

| $t_{L, L L}^{S=0, T=1}$ | $:$ | ${ }^{1} S_{0},{ }^{1} D_{2},{ }^{1} G_{4},{ }^{1} I_{6},{ }^{1} K_{8}$ |
| :---: | :---: | :---: |
| $t_{L-1, L L}^{S=1, T=1}$ | $:$ | ${ }^{3} P_{0},{ }^{3} F_{2},{ }^{3} H_{4},{ }^{3} J_{6},{ }^{3} L_{8}$ |
| $t_{L, L L}^{S=1, T=1}$ | $:$ | ${ }^{3} P_{1},{ }^{3} F_{3},{ }^{3} H_{5},{ }^{3} J_{7}$ |
| $t_{L+1, L L}^{S=1, T=1}$ | $:$ | ${ }^{3} P_{2},{ }^{3} F_{4},{ }^{3} H_{6},{ }^{3} J_{8}$ |
| $t_{L, L L}^{S, 0, T=0}$ | $:$ | ${ }^{1} P_{1},{ }^{1} F_{3},{ }^{1} H_{5},{ }^{1} J_{7}$ |
| $t_{L-1, L L}^{S=1, T=0}$ | $:$ | ${ }^{3} D_{1},{ }^{3} G_{3},{ }^{3} I_{5},{ }^{3} K_{7}$ |
| $t_{L, L L}^{S=1, T=0}$ | $:$ | ${ }^{3} D_{2},{ }^{3} G_{4},{ }^{3} I_{6},{ }^{3} K_{8}$ |
| $t_{L+1, T L=0}^{S+1, L L}$ | $:$ | ${ }^{3} S_{1},{ }^{3} D_{3},{ }^{3} G_{5},{ }^{3} I_{7}$ |

and, similarly, for the spin-orbit part:

$$
\begin{align*}
t_{p p}^{l s}= & -\frac{1}{2 \pi^{2}} \sum_{L=1}^{\infty} \frac{d P_{L}(\cos \phi)}{d \cos \phi} \frac{1}{\kappa^{\prime} \kappa}\left[-\frac{2 L-1}{L} t_{L-1, L L}^{S=1, T=1}\right. \\
& \left.-\frac{2 L+1}{L(L+1)} t_{L, L L}^{S=1, T=1}+\frac{2 L+3}{L+1} t_{L+1, L L}^{S=1, T=1}\right]  \tag{51}\\
t_{p n}^{l s}= & -\frac{1}{4 \pi^{2}} \sum_{L=1}^{\infty} \frac{d P_{L}(\cos \phi)}{d \cos \phi} \frac{1}{\kappa^{\prime} \kappa}\left[-\frac{2 L-1}{L} t_{L-1, L L}^{S=1, T=0}\right. \\
& -\frac{2 L+1}{L(L+1)} t_{L, L L}^{S=1, T=0}+\frac{2 L+3}{L+1} t_{L+1, L L}^{S=1, T=0} \\
& -\frac{2 L-1}{L} t_{L-1, L L}^{S=1, T=1}-\frac{2 L+1}{L(L+1)} t_{L, L L}^{S=1, T=1} \\
& \left.+\frac{2 L+3}{L+1} t_{L+1, L L}^{S=1, T=1}\right] . \tag{52}
\end{align*}
$$

The partial-wave components $t_{J L L}^{S T}\left(\kappa^{\prime}, \kappa ; \omega\right)$ are computed in the $N N$ center-of-mass frame, from the $N N$ potential. In this work we use two different versions of the chiral potential at the fourth order ( $\mathrm{N}^{3} \mathrm{LO}$ ) developed by Entem and Machleidt [39] and Epelbaum, Glöckle, and Meißner [44]. The $t_{J L L}^{S T}\left(\kappa^{\prime}, \kappa ; \omega\right)$ matrices are computed for each partial wave up to $J=8$. The partial waves are collected in the Table I.

## C. Transition amplitude in partial-wave representation

The optimally factorized first-order KMT optical potential as an operator in the spin space of the projectile is given in Eq. (30) as

$$
\begin{equation*}
\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)=\hat{U}^{c}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)+\frac{i}{2} \boldsymbol{\sigma} \cdot \boldsymbol{k}^{\prime} \times \boldsymbol{k} \hat{U}^{l s}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right) . \tag{53}
\end{equation*}
$$

From the conservation of the total angular momentum and parity, this spin operator can be expanded as

$$
\begin{equation*}
\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)=\frac{2}{\pi} \sum_{J L M} y_{J M}^{L \frac{1}{2}}\left(\hat{\boldsymbol{k}}^{\prime}\right) \hat{U}_{L J}\left(k^{\prime}, k ; \omega\right) y_{J M}^{L \frac{1}{2} \dagger}(\hat{\boldsymbol{k}}), \tag{54}
\end{equation*}
$$

where $J=L \pm 1 / 2$ and $y_{J M}^{L \frac{1}{2}}$ is the standard spin-angular function of Eq. (43). Inserting the expansion in Eq. (54) into

Eq. (24), we obtain the same decomposition for the $T$ matrix:

$$
\begin{equation*}
\hat{T}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right)=\frac{2}{\pi} \sum_{J L M} y_{J M}^{L \frac{1}{2}}\left(\hat{\boldsymbol{k}}^{\prime}\right) \hat{T}_{L J}\left(k^{\prime}, k ; E\right) y_{J M}^{L \frac{1}{2} \dagger}(\hat{\boldsymbol{k}}) \tag{55}
\end{equation*}
$$

where the partial-wave components of the transition operator for the elastic scattering are given by

$$
\begin{align*}
\hat{T}_{L J}\left(k^{\prime}, k ; E\right)= & \hat{U}_{L J}\left(k^{\prime}, k ; \omega\right) \\
& +\frac{2}{\pi} \int_{0}^{\infty} d p p^{2} \frac{\hat{U}_{L J}\left(k^{\prime}, p ; \omega\right) \hat{T}_{L J}(p, k ; E)}{E\left(k_{0}\right)-E(p)+i \epsilon} \tag{56}
\end{align*}
$$

where

$$
\begin{align*}
& E\left(k_{0}\right)=\sqrt{k_{0}^{2}+m_{\mathrm{proj}}^{2}}+\sqrt{k_{0}^{2}+m_{\mathrm{targ}}^{2}}  \tag{57}\\
& E(p)=\sqrt{p^{2}+m_{\mathrm{proj}}^{2}}+\sqrt{p^{2}+m_{\mathrm{targ}}^{2}} \tag{58}
\end{align*}
$$

and $m_{\text {proj }}$ and $m_{\text {targ }}$ are the masses of the projectile and of the target, respectively. In terms of the partial-wave components of the quantities $\hat{U}^{c}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ and $\hat{U}^{l s}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$, we have

$$
\begin{equation*}
\hat{U}_{L J}\left(k^{\prime}, k ; \omega\right)=\hat{U}_{L}^{c}\left(k^{\prime}, k ; \omega\right)+C_{L J} \hat{V}_{L}^{l s}\left(k^{\prime}, k ; \omega\right) \tag{59}
\end{equation*}
$$

where

$$
\begin{align*}
C_{L J} & =\frac{1}{2}\left[J(J+1)-L(L+1)-\frac{3}{4}\right] \\
\hat{V}_{L}^{l s}\left(k^{\prime}, k ; \omega\right) & =\frac{k^{\prime} k}{2 L+1}\left[\hat{U}_{L+1}^{l s}\left(k^{\prime}, k ; \omega\right)-\hat{U}_{L-1}^{l s}\left(k^{\prime}, k ; \omega\right)\right] \tag{60}
\end{align*}
$$

To obtain these results, the quantities $\hat{U}^{c}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ and $\hat{U}^{l s}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ are expanded in a manner similar to Eq. (54), with the difference that the partial-wave components $\hat{U}_{L}^{c}$ and $\hat{U}_{L}^{l s}$ are independent of $J$.

The partial-wave components of $\hat{U}^{c}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ and $\hat{U}^{l s}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ can be calculated in terms of the $N N t$-matrix components and of the nuclear densities from Eqs. (31) and (32). The projection can be performed numerically by evaluating the integral ( $a=c, l s$ )

$$
\begin{align*}
\hat{U}_{L}^{a}\left(k^{\prime}, k ; \omega\right) & =\pi^{2} \int_{-1}^{+1} d x P_{L}(x) \hat{U}^{a}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right) \\
& =\pi^{2} \int_{-1}^{+1} d x P_{L}(x) \hat{U}^{a}\left(k^{\prime}, k, x ; \omega\right) \tag{61}
\end{align*}
$$

where $x=\cos \theta$ and the potentials in terms of $k^{\prime}$ and $k$ are obtained from Eqs. (31) and (32) with

$$
\begin{align*}
q(x) & =\sqrt{k^{\prime 2}+k^{2}-2 k^{\prime} k x} \\
K(x) & =\frac{1}{2} \sqrt{k^{\prime 2}+k^{2}+2 k^{\prime} k x}  \tag{62}\\
\boldsymbol{q} \cdot \boldsymbol{K} & =\frac{1}{2}\left(k^{\prime 2}-k^{2}\right)
\end{align*}
$$

The one-dimensional integral equation for the partial-wave elements $\hat{T}_{L J}$ [Eq. (56)] is solved for the complex potentials $\hat{U}_{L J}$. In actual calculations, the number of $L$ values needed to represent the nuclear optical potential at the level of accuracy required through the partial-wave components $\hat{U}_{L J}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ can be as large as 30 for an ${ }^{16} \mathrm{O}$ target at 200 MeV .

## D. Scattering observables

Under the assumptions of parity conservation and rotational invariance, the most general form of the full amplitude for the elastic proton scattering from a spin 0 nucleus is given by

$$
\begin{equation*}
M\left(k_{0}, \theta\right)=A\left(k_{0}, \theta\right)+\sigma \cdot \hat{\boldsymbol{N}} C\left(k_{0}, \theta\right) \tag{63}
\end{equation*}
$$

where the amplitudes $A\left(k_{0}, \theta\right)$ and $C\left(k_{0}, \theta\right)$ are obtained from the partial-wave solutions of Eq. (56) as

$$
\begin{align*}
& A(\theta)=\frac{1}{2 \pi^{2}} \sum_{L=0}^{\infty}\left[(L+1) F_{L}^{+}\left(k_{0}\right)+L F_{L}^{-}\left(k_{0}\right)\right] P_{L}(\cos \theta),  \tag{64}\\
& C(\theta)=\frac{i}{2 \pi^{2}} \sum_{L=1}^{\infty}\left[F_{L}^{+}\left(k_{0}\right)-F_{L}^{-}\left(k_{0}\right)\right] P_{L}^{1}(\cos \theta) \tag{65}
\end{align*}
$$

In Eqs. (64) and (65), an implicit dependence on $k_{0}$ is assumed. The functions $F_{L}^{ \pm}$denote $F_{L J}$ for $J=L \pm 1 / 2$, respectively, and are given as

$$
\begin{equation*}
F_{L J}\left(k_{0}\right)=-\frac{A}{A-1} 4 \pi^{2} \mu\left(k_{0}\right) \hat{T}_{L J}\left(k_{0}, k_{0} ; E\right) \tag{66}
\end{equation*}
$$

where the relativistic reduced mass is

$$
\begin{equation*}
\mu\left(k_{0}\right)=\frac{E_{\text {proj }}\left(k_{0}\right) E_{\text {targ }}\left(k_{0}\right)}{E_{\text {proj }}\left(k_{0}\right)+E_{\text {targ }}\left(k_{0}\right)} \tag{67}
\end{equation*}
$$

Three independent scattering observables can be considered: the unpolarized differential cross section, the analyzing power $A_{y}$, and the spin rotation $Q$. Their expressions as functions of the amplitudes $A$ and $C$ are

$$
\begin{align*}
\frac{d \sigma}{d \Omega}(\theta) & =|A(\theta)|^{2}+|C(\theta)|^{2}  \tag{68}\\
A_{y}(\theta) & =\frac{2 \operatorname{Re}\left[A^{*}(\theta) C(\theta)\right]}{|A(\theta)|^{2}+|C(\theta)|^{2}}  \tag{69}\\
Q(\theta) & =\frac{2 \operatorname{Im}\left[A(\theta) C^{*}(\theta)\right]}{|A(\theta)|^{2}+|C(\theta)|^{2}} \tag{70}
\end{align*}
$$

## E. Treatment of Coulomb potential

In this section we include in the theoretical framework the Coulomb interaction between the incoming proton with charge $e$ and the spin-zero target with charge Ze. This has been done following the algorithm outlined in Refs. [26,27]. The interaction is separated into the sum of two parts: the "point" Coulomb interaction and the short-ranged one which is given by the sum of the nuclear potential and the short-range Coulomb interaction due to the finite dimension of the nucleus. Since the Coulomb $T$ matrix is known analytically, we only need to compute the transition matrix modified by the residual Coulomb field.

In this approach the total scattering amplitude can be written in the standard way as

$$
\begin{equation*}
M\left(k_{0}, \theta\right)=A\left(k_{0}, \theta\right)+\sigma \cdot \hat{\boldsymbol{N}} C\left(k_{0}, \theta\right) \tag{71}
\end{equation*}
$$

where now instead of Eqs. (64) and (65) we have

$$
\begin{align*}
A\left(k_{0}, \theta\right)= & F_{p t}^{c}\left(k_{0}, \theta\right)+\frac{1}{2 \pi^{2}} \sum_{L=0}^{\infty} e^{2 i \sigma_{L}} \\
& \times\left[(L+1) \bar{F}_{L}^{+}\left(k_{0}\right)+L \bar{F}_{L}^{-}\left(k_{0}\right)\right] P_{L}(\cos \theta) \tag{72}
\end{align*}
$$

$$
\begin{equation*}
C\left(k_{0}, \theta\right)=\frac{i}{2 \pi^{2}} \sum_{L=1}^{\infty} e^{2 i \sigma_{L}}\left[\bar{F}_{L}^{+}\left(k_{0}\right)-\bar{F}_{L}^{-}\left(k_{0}\right)\right] P_{L}^{1}(\cos \theta) \tag{73}
\end{equation*}
$$

In Eqs. (72) and (73) $F_{p t}^{c}\left(k_{0}, \theta\right)$ is the Coulomb scattering amplitude due to a point charge [61]:

$$
\begin{equation*}
F_{p t}^{c}\left(k_{0}, \theta\right)=\frac{-\eta\left(k_{0}\right) 2^{i \eta\left(k_{0}\right)} \exp \left[2 i \sigma_{0}-i \eta\left(k_{0}\right) \ln (1-\cos \theta)\right]}{k_{0}(1-\cos \theta)} \tag{74}
\end{equation*}
$$

where

$$
\begin{equation*}
\eta(k)=\frac{\mu Z \alpha}{k} \tag{75}
\end{equation*}
$$

is the Sommerfeld parameter, $\mu$ is the reduced mass of Eq. (67), and $\alpha$ is the fine structure constant. The Coulomb phase shifts $\sigma_{L}$ are given by

$$
\begin{equation*}
\sigma_{L}=\arg \Gamma\left[L+1+i \eta\left(k_{0}\right)\right] . \tag{76}
\end{equation*}
$$

The partial-wave scattering amplitudes $\bar{F}_{L}^{ \pm}$are obtained from the solution of the Coulomb-distorted $\bar{T}$ matrix

$$
\begin{equation*}
\bar{T}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right)=\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)+\int d^{3} p \frac{\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{p} ; \omega\right) \bar{T}(\boldsymbol{p}, \boldsymbol{k} ; E)}{E\left(k_{0}\right)-E(p)+i \epsilon}, \tag{77}
\end{equation*}
$$

where

$$
\begin{equation*}
\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)=\left\langle\boldsymbol{k}^{\prime}\right| \bar{U}(\omega)|\boldsymbol{k}\rangle=\left\langle\psi_{c}^{(+)}\left(\boldsymbol{k}^{\prime}\right)\right| \hat{U}(\omega)\left|\psi_{c}^{(+)}(\boldsymbol{k})\right\rangle \tag{78}
\end{equation*}
$$

and $\psi_{c}^{(+)}(\boldsymbol{k})$ is the Coulomb-distorted wave function.
In order to solve Eq. (77), we need to be able to generate the momentum-space matrix element $\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ as given in Eq. (78). We begin with the potential $\hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$, discussed in Sec. II A, and we transform it into the coordinate space through the double Fourier transform

$$
\begin{equation*}
\hat{U}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r} ; \omega\right)=\int d^{3} k^{\prime} d^{3} k\left\langle\boldsymbol{r}^{\prime} \mid \boldsymbol{k}^{\prime}\right\rangle \hat{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)\langle\boldsymbol{k} \mid \boldsymbol{r}\rangle \tag{79}
\end{equation*}
$$

and then we construct the matrix element of Eq. (78) by folding $\hat{U}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r} ; \omega\right)$ with coordinate space Coulomb wave functions

$$
\begin{align*}
\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)= & \int d^{3} r^{\prime} d^{3} r\left\langle\psi_{c}^{(+)}\left(\boldsymbol{k}^{\prime}\right) \mid \boldsymbol{r}^{\prime}\right\rangle \\
& \times \hat{U}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r} ; \omega\right)\left\langle\boldsymbol{r} \mid \psi_{c}^{(+)}(\boldsymbol{k})\right\rangle \tag{80}
\end{align*}
$$

In the partial-wave representation, Eq. (79) for the central and spin-orbit parts becomes

$$
\begin{align*}
\hat{U}_{L}^{a}\left(r^{\prime}, r ; \omega\right)= & \frac{4}{\pi^{2}} \int_{0}^{\infty} d k^{\prime} k^{\prime 2} \int_{0}^{\infty} d k k^{2} j_{L}\left(k^{\prime} r^{\prime}\right) \\
& \times \hat{U}_{L}^{a}\left(k^{\prime}, k ; \omega\right) j_{L}(k r) \tag{81}
\end{align*}
$$

where $j_{L}(k r)$ are spherical Bessel functions. Similarly, Eq. (80) becomes

$$
\begin{align*}
\bar{U}_{L}^{a}\left(k^{\prime}, k ; \omega\right)= & \frac{1}{k^{\prime} k} \int_{0}^{\infty} d r^{\prime} r^{\prime} \int_{0}^{\infty} d r r F_{L}\left(\eta, k^{\prime} r^{\prime}\right) \\
& \times \hat{U}_{L}^{a}\left(r^{\prime}, r ; \omega\right) F_{L}(\eta, k r) \tag{82}
\end{align*}
$$

where $F_{L}$ is the regular Coulomb function. The potential $\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)$ can be expanded in partial waves as in Eq. (54):

$$
\bar{U}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; \omega\right)=\frac{2}{\pi} \sum_{J L M} y_{J M}^{L \frac{1}{2}}\left(\hat{\boldsymbol{k}}^{\prime}\right) \bar{U}_{L J}\left(k^{\prime}, k ; \omega\right) y_{J M}^{L \frac{1}{2} \dagger}(\hat{\boldsymbol{k}}),
$$

where

$$
\begin{equation*}
\bar{U}_{L J}\left(k^{\prime}, k ; \omega\right)=\bar{U}_{L}^{c}\left(k^{\prime}, k ; \omega\right)+C_{L J} \bar{V}_{L}^{l s}\left(k^{\prime}, k ; \omega\right), \tag{84}
\end{equation*}
$$

and

$$
\begin{align*}
C_{L J} & =\frac{1}{2}\left[J(J+1)-L(L+1)-\frac{3}{4}\right] \\
\bar{V}_{L}^{l s}\left(k^{\prime}, k ; \omega\right) & =\frac{k^{\prime} k}{2 L+1}\left[\bar{U}_{L+1}^{l s}\left(k^{\prime}, k ; \omega\right)-\bar{U}_{L-1}^{l s}\left(k^{\prime}, k ; \omega\right)\right] \tag{85}
\end{align*}
$$

Likewise, we can expand the $\bar{T}$ matrix in Eq. (77) as

$$
\begin{equation*}
\bar{T}\left(\boldsymbol{k}^{\prime}, \boldsymbol{k} ; E\right)=\frac{2}{\pi} \sum_{J L M} y_{J M}^{L \frac{1}{2}}\left(\hat{\boldsymbol{k}}^{\prime}\right) \bar{T}_{L J}\left(k^{\prime}, k ; E\right) y_{J M}^{L \frac{1}{2} \dagger}(\hat{\boldsymbol{k}}), \tag{86}
\end{equation*}
$$

where the partial-wave components are

$$
\begin{align*}
\bar{T}_{L J}\left(k^{\prime}, k ; E\right)= & \bar{U}_{L J}\left(k^{\prime}, k ; \omega\right) \\
& +\frac{2}{\pi} \int_{0}^{\infty} d p p^{2} \frac{\bar{U}_{L J}\left(k^{\prime}, p ; \omega\right) \bar{T}_{L J}(p, k ; E)}{E\left(k_{0}\right)-E(p)+i \epsilon} . \tag{87}
\end{align*}
$$

The partial-wave scattering amplitudes $\bar{F}_{L}^{ \pm}$entering Eqs. (72) and (73) are given by

$$
\begin{equation*}
\bar{F}_{L J}\left(k_{0}\right)=-\frac{A}{A-1} 4 \pi^{2} \mu\left(k_{0}\right) \bar{T}_{L J}\left(k_{0}, k_{0} ; E\right) \tag{88}
\end{equation*}
$$

## III. THE NN AMPLITUDES

In this section we present and discuss the theoretical results for the $p p$ and $p n$ Wolfenstein amplitudes which are used to compute the central $a(45)$ and the spin-orbit part $c(46)$ of the three-dimensional $N N t$ matrix. Calculations are performed by using two different versions of the chiral potential at fourth order ( $\mathrm{N}^{3} \mathrm{LO}$ ) based on the works of Entem and Machleidt [39-41,43] and Epelbaum et al. [44]. The performance of our code has been tested against the CD-Bonn potential [60] reproducing well-known results $[63,64]$ in order to check its numerical correctness.

Entem and Machleidt (EM), who first presented a chiral potential at the fourth order, treat divergent terms in the two-pion exchange (2PE) contributions with dimensional regularization (DR), while Epelbaum, Glöckle, and Meißner (EGM) employ a spectral function regularization (SFR). In both cases the goal is to cut out the short-range part of the 2PE contribution that, as shown in Ref. [40], has unphysically strong attraction, particularly at $\mathrm{N}^{2} \mathrm{LO}$ (for a comprehensive discussion about different regularization schemes we refer the reader to Sec. 3.2.1 of Ref. [44]). As a usual procedure, the nucleon-nucleon potential entering the LS equation is multiplied by a regulator function $f^{\Lambda}$ :

$$
\begin{equation*}
V\left(\boldsymbol{k}, \boldsymbol{k}^{\prime}\right) \rightarrow V\left(\boldsymbol{k}, \boldsymbol{k}^{\prime}\right) f^{\Lambda}\left(k, k^{\prime}\right) \tag{89}
\end{equation*}
$$

where

$$
\begin{equation*}
f^{\Lambda}=\exp \left[-\left(k^{\prime} / \Lambda\right)^{2 n}-(k / \Lambda)^{2 n}\right] \text { with } n=2,3 . \tag{90}
\end{equation*}
$$

While Entem and Machleidt present results for three choices of the cutoff necessary to regulate the high-momentum components in the LS equation ( $\Lambda=450,500$, and 600 MeV ),


FIG. 1. Real (left panel) and imaginary (right panel) parts of $p p$ and $p n$ Wolfenstein amplitudes ( $a$ and $c$ ) as functions of the center-of-mass $N N$ angle $\phi$. All the amplitudes are computed at 100 MeV using the EM potentials [39-41,43] with a LS cutoff ranging from 450 to 600 MeV . Data (black squares) are taken from Ref. [62].

Epelbaum et al. [44] allow us also to study variations of the cutoff $\tilde{\Lambda}$ that regulates the 2 PE contribution. In fact, in the latter approach one can choose between the following cutoff combinations:

$$
\begin{align*}
\{\Lambda, \tilde{\Lambda}\}= & \{450,500\},\{450,700\},\{550,600\} \\
& \{600,600\},\{600,700\} \tag{91}
\end{align*}
$$

In the following figures all the results are labeled by an acronym (to distinguish the authors) followed by the value of the LS cutoff ( $\Lambda$ ). In the EGM case, for $\Lambda=450$ and 600 MeV we plot bands that show how calculations can change with respect to variations of the SFR cutoff $\tilde{\Lambda}$.

In Fig. 1 the theoretical results for the real and imaginary parts of $p p$ and $p n$ Wolfenstein amplitudes ( $a$ and $c$ ) computed at an energy of 100 MeV are shown as functions of the center-of-mass $N N$ angle $\phi$ and compared with the experimental data. The calculations are performed by using the EM potentials [39-41,43] with a LS cutoff ranging from 450 to 600 MeV . The experimental data are globally reproduced by the three potentials, with the only remarkable exception of the real part of the $c_{p p}$ amplitude that is overestimated. It must be considered, however, that this is a small quantity, i.e., two orders of magnitude smaller than the respective imaginary part, and it will only provide a very small contribution to the optical potential. Concerning the other amplitudes, some deviations from the experimental data are found, in particular for the imaginary part of the $c_{p p}$ and $c_{p n}$ amplitudes. Finding some discrepancies is not surprising, because the NN amplitudes are directly related to the empirical $N N$ phase shifts, which are not always perfectly reproduced by realistic potentials for some $J$ (see ${ }^{3} F_{3},{ }^{3} F_{4}$, and ${ }^{3} G_{5}$ cases in Figs. 8 and 9 of Ref. [39] and in Fig. 27 of Ref. [44]).

In Fig. 2 we show the results obtained at 100 MeV with the EGM potentials [44]. Also in this case, all three potentials are


FIG. 2. The same as in Fig. 1 using EGM potentials [44] with a LS cutoff ranging from 450 to 600 MeV . In two cases ( $\Lambda=450$ and 600 MeV ) we show uncertainty bands produced by changing $\tilde{\Lambda}$ according to Eq. (91). Data (black squares) are taken from Ref. [62].
in overall good agreement with the experimental data with the only remarkable exception of the real part of the $c_{p p}$ amplitude. In particular, they show very similar results and in many cases the yellow and turquoise bands are overlapped. Their trends are also very close to the ones shown in Fig. 1 for the EM potential and they display the same discrepancy in comparison with the experimental data for $c_{p n}$ and around the peak of the imaginary part of $c_{p p}$.

Since ChPT is a low-momentum expansion of QCD, we expect that, as the energy is increased, larger discrepancies appear with respect to empirical data. In Figs. 3 and 4 we present the results corresponding to those shown in Figs. 1 and 2 but at an energy of 200 MeV . As energy is increased, all potentials are still unable to reproduce the experimental data


FIG. 3. Same as in Fig. 1 but for an energy of 200 MeV . Data (black squares) are taken from Ref. [62].


FIG. 4. Same as is Fig. 2 but for an energy of 200 MeV . Data (black squares) are taken from Ref. [62].
of the real part of the $c_{p p}$ amplitude, but most of the chiral potentials give satisfactory results, in agreement with the data for all the other amplitudes, with one notable exception. In fact, in both approaches, potentials with a cutoff of 450 MeV (see Figs. 3 and 4) fail to reproduce the real part of $a_{p p}$ and $a_{p n}$ and underestimate the imaginary part of the $c_{p n}$ amplitude. Based on these flaws, we predict an unsatisfactory result for a nucleon-nucleus optical potential if EM-450 or EGM-450 are employed at energies well above 100 MeV .

Chiral potentials at orders NLO, $\mathrm{N}^{2} \mathrm{LO}$, and $\mathrm{N}^{3} \mathrm{LO}$ have been constructed and compared in Ref. [44] for EGM and in Ref. [42] for EM. As an example, in Fig. 5 we present a


FIG. 5. Real (left panel) and imaginary (right panel) parts of $p p$ and $p n$ Wolfenstein amplitudes ( $a$ and $c$ ) as functions of the center-of-mass $N N$ angle $\phi$ computed at 200 MeV using the EGM potential [44] at different orders: red bands are NLO results, green and blue bands are, respectively, the $\mathrm{N}^{2} \mathrm{LO}$ and $\mathrm{N}^{3} \mathrm{LO}$ results. Data (black squares) are taken from Ref. [62].
systematic study order by order of the convergence pattern using the EGM potential. At each order, calculations have been performed for all the cutoff combinations in Eq. (91), the variation produced by the different combinations in the calculated amplitudes is depicted by the bands in the figure. From the results shown for all $N N$ amplitudes, we can draw the conclusion that it is mandatory to use potentials at order $\mathrm{N}^{3} \mathrm{LO}$. At orders NLO and $\mathrm{N}^{2} \mathrm{LO}$ the amplitudes not only underestimate or overestimate empirical data but also miss the overall shapes. The order-by-order convergence will be further explored in a forthcoming presentation using the recent $\mathrm{N}^{4} \mathrm{LO}$ potential [45].

## IV. SCATTERING RESULTS

In this section we present and discuss our numerical results for the $N A$ elastic scattering observables calculated with the microscopic optical potential obtained within the theoretical framework described in Sec. II. As a study case in our calculations we consider elastic proton scattering on ${ }^{16} \mathrm{O}$.

We investigate the sensitivity of our results to the choice of the $N N$ potential and, in particular, their dependence on the cutoff values. In order to investigate and emphasize the differences between the different $N N$ potentials and also on the basis of the results obtained for the $N N$ Wolfenstein amplitudes $a$ and $c$, the scattering observables have been calculated for different energies $(100,135,200$, and 318 MeV ) for which experimental data are available. In light of the fact that chiral potentials are based on a low-momentum expansion, the last energy may be considered beyond the limit of applicability of such potentials.

With these calculations we intend to achieve the following goals: (1) to check the agreement of our theoretical predictions with the empirical data, (2) to study the limits of applicability of chiral potentials in terms of the proton energy, (3) to identify the best set of values for the LS and, eventually, SFR cutoffs.

In Figs. 6-9 we show the differential cross section $(d \sigma / d \Omega)$, the analyzing power $A_{y}$, and the spin rotation $Q$ for elastic proton scattering on ${ }^{16} \mathrm{O}$ as functions of the center-of-mass scattering angle $\theta$ with the above-mentioned four energies ( $E=100,135,200$, and 318 MeV ). In the left panels we show the results obtained with the EM potentials [39-41,43] while in the right panels we show the results obtained with the EGM potentials [44]. All potentials are denoted by the value of the LS cutoff. The Coulomb interaction between the proton and the target nucleus is included as described in Sec. II E.

In Fig. 6, at 100 MeV , all sets of potentials, regardless of cutoffs and theoretical approaches, give very similar results for all three observables, with the exception of $A_{y}$ above 50 degrees, where all potentials overestimate the experimental data up to the maximum and then display an unrealistic downward trend, and $Q$ around the maximum at 30 degrees. In particular, the experimental cross section is well reproduced by all potentials in the minimum region, between 30 and 35 degrees. Polarization observables are usually more sensitive to the differences in the potentials and to the ingredients and approximations of the model. Experimental data for such observables are usually more difficult to reproduce. Even if


FIG. 6. Scattering observables (differential cross section $d \sigma / d \Omega$, analyzing power $A_{y}$, and spin rotation $Q$ ) as a function of the center-of-mass scattering angle $\theta$ for elastic proton scattering on ${ }^{16} \mathrm{O}$ computed at 100 MeV (laboratory energy). On the left panel we employ the set of EM potentials [39-41,43] while in the right panel we show the EGM potentials [44]. All potentials are denoted by the value of the LS cutoff. Coulomb distortion is included as explained in Sec. II E. Data are taken from Refs. [65,66].
differences are rather small, potentials with the largest cutoff ( $\Lambda=600 \mathrm{MeV}$ ) seem to provide the best description of $A_{y}$.

A similar result is obtained in Fig. 7, where we display the scattering observables calculated at 135 MeV . In this case all sets of potentials reproduce very well the experimental cross section and globally describe the shape of $A_{y}$ but are unable to reproduce its magnitude for angles larger than $20^{\circ}$.

In Fig. 8 we plot the results obtained at 200 MeV . At this energy, it is clear that potentials obtained with the lower cutoffs (EM-450 and EGM-450) cannot be employed any further: in both cases, the differential cross sections are not satisfactorily reproduced and the behavior of $A_{y}$ and $Q$ as a function of $\theta$ is in clear disagreement with the empirical one. On the other hand, the remaining sets of potentials well describe the experimental cross sections and the analyzing power $A_{y}$, which is reasonably described not only for small scattering angles but also for values larger than the minimum value up to about $45^{\circ}$.


FIG. 7. The same as is Fig. 6 but for an energy of 135 MeV . Data are taken from Refs. [65,66].


FIG. 8. The same as is Fig. 6 but for an energy of 200 MeV . Data are taken from Refs. [65,66].

On the basis of all these results for ${ }^{16} \mathrm{O}$ we can draw two conclusions: (1) Potentials with lower cutoffs cannot reproduce experimental data at energies close to 200 MeV . (2) There is no appreciable difference in using 500 or 600 MeV as LS cutoffs, even if the EM-600 and EGM-600 potentials seem to have a slightly better agreement with empirical data; in particular looking at polarization observables.

For energies above 200 MeV , this behavior changes and the agreement with the experimental data begins to fail. This failure becomes larger as the energy increases. As an example, in Fig. 9 we display the results for the scattering observables on ${ }^{16} \mathrm{O}$ computed at 318 MeV , an energy for which experimental data are available. We clearly see that, at this energy, all potentials are unable to describe the data. A somewhat better description is given by the EM-600 potential, which is able to reproduce the global shape of the experimental results and the position of the minima, but the general agreement is poor. However, we stress that ChPT is a low-momentum expansion and its goal should be to perform calculations at lower energies.

In Fig. 10 we repeat the same order by order analysis (NLO, $\mathrm{N}^{2} \mathrm{LO}$, and $\mathrm{N}^{3} \mathrm{LO}$ ) of the convergence pattern performed in Sec. III for the $N N$ amplitudes. The results confirm the


FIG. 9. The same as is Fig. 6 but for an energy of 318 MeV . Data are taken from Refs. [65,66].


FIG. 10. Scattering observables as a function of the center-ofmass scattering angle $\theta$ for elastic proton scattering on ${ }^{16} \mathrm{O}$ computed at 200 MeV (laboratory energy) with the EGM potential [44] at different orders: red bands are the NLO results, green and blue bands are respectively the $\mathrm{N}^{2} \mathrm{LO}$ and $\mathrm{N}^{3} \mathrm{LO}$ results. Data are taken from Refs. [65,66].
conclusion drawn looking at the $N N$ amplitudes, i.e., that it is mandatory to use potentials at order $\mathrm{N}^{3} \mathrm{LO}$. At orders NLO and $\mathrm{N}^{2} \mathrm{LO}$ our theoretical predictions not only underestimate or overestimate empirical data but also miss the overall shapes. The order-by-order convergence suggests that there is space for improvement going to higher orders $\left(\mathrm{N}^{4} \mathrm{LO}\right)$ [45].

In order to understand why some potentials provide a better description of certain scattering data than other potentials, in Figs. 11 and 12 we plot the relevant components, Eqs. (64) and


FIG. 11. Contributions, in terms of the amplitudes $A$ (64) and $C$ (65), to the scattering observables (differential cross section $d \sigma / d \Omega$ and asymmetry parameter $A_{y}$ ) for elastic proton scattering on ${ }^{16} \mathrm{O}$ computed at 200 MeV (laboratory energy) using two EGM potentials with $\{\Lambda, \tilde{\Lambda}\}=\{600,600\},\{600,700\}$. Coulomb distortion is included. Data are taken from Refs. [65,66].


FIG. 12. $L$ components (92) for the $C$ amplitude for two EGM potentials with the following cutoffs: $\{\Lambda, \tilde{\Lambda}\}=\{600,600\},\{600,700\}$.
(65), for the differential cross section and the analyzing power computed at 200 MeV . We have chosen, as a test case, two EGM potentials, with $\{\Lambda, \tilde{\Lambda}\}=\{600,600\},\{600,700\}$, that reproduce differential cross sections with the same accuracy but give different predictions for the analyzing power. In the upper panels of Fig. 11 we plot, for both potentials, the total differential cross section (proportional to the sum $|A|^{2}+|C|^{2}$ ) with a red line, and the single contributions $|A|^{2}$ and $|C|^{2}$ with green and blue lines, respectively. The two potentials give similar results for $|A|^{2}$ while significant differences around the minima are obtained for $|C|^{2}$. These differences, however, do not affect the final cross section, which is clearly dominated by the contribution proportional to $|A|^{2}$. The two potentials give relevant differences for the analyzing power $A_{y}$, which is plotted in the lower panels. In this case we cannot disentangle single contributions because $A_{y}$ is proportional to a combination of $A$ and $C\left(A_{y} \sim \operatorname{Re}\left[A^{*}(\theta) C(\theta)\right]\right)$. Nonetheless, a connection between $|C|^{2}$ in the upper blue curves and $A_{y}$ seems to be plausible. To test if the first minimum of $A_{y}$ is really determined by the behavior of the $C$ amplitudes, in Fig. 12 we plot the $L$ components of $C$, defined as

$$
\begin{equation*}
C_{L}=\left[F_{L}^{+}\left(k_{0}\right)-F_{L}^{-}\left(k_{0}\right)\right] P_{L}^{1}(\cos \theta) \tag{92}
\end{equation*}
$$

evaluated at the angle $\theta=27^{\circ}$ corresponding to the minimum position. The two potentials give close results for the real parts and large differences for the imaginary parts of the $L$ components. With both potentials the real part of the $C$ amplitude is almost canceled in the sum $C(\theta)=i /\left(2 \pi^{2}\right) \sum_{L} C_{L}$. For the imaginary part, the sum gives a contribution that is small for $\{\Lambda, \tilde{\Lambda}\}=\{600,600\}$ and sizable for $\{\Lambda, \tilde{\Lambda}\}=\{600,700\}$. As a consequence, in the case $\{\Lambda, \tilde{\Lambda}\}=\{600,600\}$ the $C$ amplitude is very small and the analyzing power and the $|C|^{2}$ contribution to the differential cross section develop well-defined minima, while in the case $\{\Lambda, \tilde{\Lambda}\}=\{600,700\}$, where the $C$ amplitude is larger, the corresponding minima are not deep enough and the disagreement with the experimental $A_{y}$ is more pronounced.

## v. CONCLUSIONS

In this work we obtained a new microscopic optical potential for elastic proton-nucleus scattering. Our optical potential was derived as the first-order term within the spectator expansion of the nonrelativistic multiple scattering theory. In the interaction between the projectile and the target nucleon, which is described by the $N N \tau$ matrix, we neglected medium effects and adopted the impulse approximation, which consists of replacing $\tau$ by the free $N N t$ matrix.

As a further simplification, we adopted the optimum factorization approximation, where the optical potential is given in a factorized form by the product of the free $N N t$ matrix and the nuclear density. This form conserves the off-shell nature of the optical potential and has been used in this work to compute the cross sections and the polarization observables of elastic proton-nucleus scattering.

Two basic ingredients underlie the calculation of our microscopic optical potential: the $N N$ interaction and a model for nuclear densities. For the $N N$ interaction we used here for the first time the chiral potential. Microscopic optical potentials have been derived from two different versions of the chiral potential at fourth order $\left(\mathrm{N}^{3} \mathrm{LO}\right)$ based on the work of Entem and Machleidt (EM) [39-41,43] and Epelbaum, Glöckle, and Meißner (EGM) [44], which differ in the regularization scheme employed in the two-pion exchange term and in the choice of the cutoffs. Neutron and proton densities have been obtained by considering a system of nucleons coupled to the exchange of mesons and the electromagnetic field through an effective Lagrangian. In practice, they were computed within the RMF description [52] of spherical nuclei by using a DDME model [52]. The Coulomb proton-nucleus interaction has also been included in the calculations.

The $N N$ potentials have been used to calculate the $N N$ amplitudes that have then been employed to compute the $N N t$ matrix. Results for $p p$ and $n p$ Wolfenstein amplitudes ( $a$ and $c$ ) obtained with different $N N$ potentials have been presented and discussed. Since ChPT is a low-momentum expansion of QCD, the agreement of the chiral potential with the experimental data becomes, as expected, worse upon increasing the energy. While at 100 MeV all the $N N$ potentials are able to reproduce the experimental amplitudes, with the only exception of the real part of $c_{p p}$ amplitude, which is anyhow extremely small, at 200 MeV the set of potentials with lower cutoffs $(450 \mathrm{MeV})$ fail to reproduce empirical data.

As a case study for our investigation we considered elastic proton scattering on ${ }^{16} \mathrm{O}$. Results for the cross section, the analyzing power, and the spin rotation have been presented and discussed in comparison with available experimental data. Calculations have been performed with different $N N$ potentials at different energies.

The comparison between the results obtained with the different versions of the chiral potential represents a useful test of the reliability of our new optical potentials and allows us to identify the best set of LS cutoff values.

Polarization observables are more sensitive to the differences in the $N N$ interactions and to the approximations of the model. This sensitivity makes it difficult to describe the experimental analyzing powers over the whole scattering angular distribution. The optical potentials obtained from all the $N N$ potentials give close results and a good description of the experimental cross sections for proton energies up to about 135 MeV . Of course, the differences among the results obtained with different $N N$ potentials increase with the energy and with the scattering angle. Our results indicate that EM-600 and EGM-600 provide a slightly better agreement with empirical data for energies up to 200 MeV . Increasing the energy, however, the agreement between the results from chiral potentials and data declines and it is plausible to believe that, above 200 MeV , ChPT is no longer applicable.

In the near future we plan to study the order-by-order convergence by using the recent $\mathrm{N}^{4} \mathrm{LO}$ potential [45] and to improve our calculations including three-body forces and nuclear medium effects. In addition, our investigation will be extended to $N \neq Z$ nuclei [67].

The case of elastic proton scattering considered in this work represents the first natural and necessary test of the reliability of an optical potential. The optical potential, however, represents a crucial and critical input for calculations over a wide variety of nuclear reactions and can therefore be employed in many other situations beyond those considered in this paper.

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